

THE UNIVERSITY OF BRITISH COLUMBIA

Physics 401 Assignment # 6:

ELECTROMAGNETIC WAVES

Wed. 8 Feb. 2006 — finish by Wed. 22 Feb.

1. **CMBR:** Most of the electromagnetic energy in the universe is in the cosmic microwave background radiation (CMBR), sometimes referred to as the 3° Kelvin background. Penzias and Wilson discovered the CMBR in 1965 using a radio telescope, and subsequently received the Nobel Prize for this discovery. This background radiation has wavelength $\lambda \sim 1.1$ mm. The energy density of the CMBR is about 4.0×10^{-14} J/m³. What is the *rms* electric field strength of the CMBR?

2. **STANDING WAVES:** Consider standing electromagnetic waves:

$$\vec{E} = E_0 (\sin kz \sin \omega t) \hat{x} \quad \text{with} \quad \vec{B} = B_0 (\cos kz \cos \omega t) \hat{y} .$$

- (a) Show that these satisfy the wave equation (9.2).
- (b) Show that we must also have $c = \omega/k$ and $E_0 = cB_0$.
- (c) Show that the time-averaged power flow across *any* area will be zero.
- (d) Show that the Poynting vector will also be zero, *i.e.* there is no net energy flow.

3. (p. 386, Problem 9.14) — **REFLECTED & TRANSMITTED POLARIZATION:** In Eqs. (9.76) and (9.77) it was tacitly assumed that the reflected and transmitted waves have the same *polarization* as the incident wave, namely along the \hat{x} direction. Prove that this *must* be so. [Hint: Let the polarization vectors of the reflected and transmitted waves be

$$\hat{n}_T = \cos \theta_T \hat{x} + \sin \theta_T \hat{y} \quad \text{and} \quad \hat{n}_R = \cos \theta_R \hat{x} + \sin \theta_R \hat{y}$$

and prove from the boundary conditions that $\theta_T = \theta_R = 0$.]

4. (p. 392, Problem 9.15) — **COMPLEX ALGEBRA EXERCISE:** Suppose that we have six nonzero constants A, B, C, a, b, c such that $Ae^{iax} + Be^{ibx} = Ce^{icx}$ for all x . Prove that $a = b = c$ and $A + B = C$.

5. (p. 392, Problem 9.17) — **DIAMOND:** The index of refraction of diamond is 2.42. Construct the graph analogous to Figure 9.16 for the air/diamond interface. (Assume $\mu_1 = \mu_2 = \mu_0$.) In particular, calculate

- (a) the amplitudes at normal incidence;
- (b) Brewster's angle; and
- (c) the "crossover" angle at which the reflected and transmitted amplitudes are equal.

6. **PLANE WAVE STRESS TENSOR:** Find all the elements of the Maxwell stress tensor of a monochromatic plane wave traveling in the z -direction, polarized in the x -direction:

$$\vec{E}(z, t) = E_0 \cos(kz - \omega t + \delta) \hat{x}$$

$$\vec{B}(z, t) = \frac{E_0}{c} \cos(kz - \omega t + \delta) \hat{y}$$

In what direction does this EM wave transport momentum? Does this agree with the form of the Maxwell stress tensor you just deduced?

7. (p. 412, Problem 9.33) — **SPHERICAL WAVES:** Suppose

$$\vec{E}(r, \theta, \phi, t) = A \frac{\sin \theta}{r} \left[\cos(kr - \omega t) - \left(\frac{1}{kr} \right) \sin(kr - \omega t) \right] \hat{\phi} \quad \text{with} \quad c = \frac{\omega}{k},$$

as usual. [This is, incidentally, the simplest possible **spherical wave**. For notational convenience, let $(kr - \omega t) \equiv u$ in your calculations.]

- (a) Show that \vec{E} obeys all four of Maxwell's equations, in vacuum, and find the associated magnetic field.
- (b) Calculate the Poynting vector. Average \vec{S} over a full cycle to get the intensity vector \vec{I} . Does \vec{I} point in the expected direction? Does it fall off like r^{-2} , as it should?
- (c) Integrate $\vec{I} \cdot d\vec{a}$ over a spherical surface to determine the total power radiated. [You should get $P = 4\pi A^2 / 3\mu_0 c$.]