

# NUCLEAR PHYSICS

an

*Introduction*

by

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# The Periodic Table of the Elements

		1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18														
												Pnictogens		Chalcogens		Halogens																	
1	<b>H</b> Hydrogen 1.008	<table border="1"> <tr> <td colspan="2">Metals</td> <td colspan="2">Metalloids</td> <td colspan="2">Nonmetals</td> </tr> <tr> <td>Alkali metals</td> <td>Alkaline earth metals</td> <td>Lanthanoids</td> <td>Transition metals</td> <td>Post-transition metals</td> <td>Reactive nonmetals</td> <td>Noble gases</td> </tr> <tr> <td></td> <td></td> <td>Actinoids</td> <td></td> <td></td> <td></td> <td></td> </tr> </table>										Metals		Metalloids		Nonmetals		Alkali metals	Alkaline earth metals	Lanthanoids	Transition metals	Post-transition metals	Reactive nonmetals	Noble gases			Actinoids					2	<b>He</b> Helium 4.0026
Metals		Metalloids		Nonmetals																													
Alkali metals	Alkaline earth metals	Lanthanoids	Transition metals	Post-transition metals	Reactive nonmetals	Noble gases																											
		Actinoids																															
2	<b>Li</b> Lithium 6.94	<b>Be</b> Beryllium 9.0122											<b>B</b> Boron 10.81	<b>C</b> Carbon 12.011	<b>N</b> Nitrogen 14.007	<b>O</b> Oxygen 15.999	<b>F</b> Fluorine 18.998	<b>Ne</b> Neon 20.180															
3	<b>Na</b> Sodium 22.990	<b>Mg</b> Magnesium 24.305											<b>Al</b> Aluminium 26.982	<b>Si</b> Silicon 28.085	<b>P</b> Phosphorus 30.974	<b>S</b> Sulfur 32.06	<b>Cl</b> Chlorine 35.45	<b>Ar</b> Argon 39.948															
4	<b>K</b> Potassium 39.098	<b>Ca</b> Calcium 40.078	<b>Sc</b> Scandium 44.956	<b>Ti</b> Titanium 47.867	<b>V</b> Vanadium 50.942	<b>Cr</b> Chromium 51.996	<b>Mn</b> Manganese 54.938	<b>Fe</b> Iron 55.845	<b>Co</b> Cobalt 58.933	<b>Ni</b> Nickel 58.693	<b>Cu</b> Copper 63.546	<b>Zn</b> Zinc 65.38	<b>Ga</b> Gallium 69.723	<b>Ge</b> Germanium 72.630	<b>As</b> Arsenic 74.922	<b>Se</b> Selenium 78.971	<b>Br</b> Bromine 79.904	<b>Kr</b> Krypton 83.798															
5	<b>Rb</b> Rubidium 85.468	<b>Sr</b> Strontium 87.62	<b>Y</b> Yttrium 88.906	<b>Zr</b> Zirconium 91.224	<b>Nb</b> Niobium 92.906	<b>Mo</b> Molybdenum 95.95	<b>Tc</b> Technetium (98)	<b>Ru</b> Ruthenium 101.07	<b>Rh</b> Rhodium 102.91	<b>Pd</b> Palladium 106.42	<b>Ag</b> Silver 107.87	<b>Cd</b> Cadmium 112.41	<b>In</b> Indium 114.82	<b>Sn</b> Tin 118.71	<b>Sb</b> Antimony 121.76	<b>Te</b> Tellurium 127.60	<b>I</b> Iodine 126.90	<b>Xe</b> Xenon 131.29															
6	<b>Cs</b> Caesium 132.91	<b>Ba</b> Barium 137.33	57-71	<b>Hf</b> Hafnium 178.49	<b>Ta</b> Tantalum 180.95	<b>W</b> Tungsten 183.84	<b>Re</b> Rhenium 186.21	<b>Os</b> Osmium 190.23	<b>Ir</b> Iridium 192.22	<b>Pt</b> Platinum 195.08	<b>Au</b> Gold 196.97	<b>Hg</b> Mercury 200.59	<b>Tl</b> Thallium 204.38	<b>Pb</b> Lead 207.2	<b>Bi</b> Bismuth 208.98	<b>Po</b> Polonium (209)	<b>At</b> Astatine (210)	<b>Rn</b> Radon (222)															
7	<b>Fr</b> Francium (223)	<b>Ra</b> Radium (226)	89-103	<b>Rf</b> Rutherfordium (267)	<b>Db</b> Dubnium (268)	<b>Sg</b> Seaborgium (269)	<b>Bh</b> Bohrium (270)	<b>Hs</b> Hassium (277)	<b>Mt</b> Meitnerium (278)	<b>Ds</b> Darmstadtium (281)	<b>Rg</b> Roentgenium (282)	<b>Cn</b> Copernicium (285)	<b>Nh</b> Nihonium (286)	<b>Fl</b> Flerovium (289)	<b>Mc</b> Moscovium (290)	<b>Lv</b> Livermorium (293)	<b>Ts</b> Tennessine (294)	<b>Og</b> Oganesson (294)															
For elements with no stable isotopes, the mass number of the isotope with the longest half-life is in parentheses.																																	
		6	<b>La</b> Lanthanum 138.91	<b>Ce</b> Cerium 140.12	<b>Pr</b> Praseodymium 140.91	<b>Nd</b> Neodymium 144.24	<b>Pm</b> Promethium (145)	<b>Sm</b> Samarium 150.36	<b>Eu</b> Europium 151.96	<b>Gd</b> Gadolinium 157.25	<b>Tb</b> Terbium 158.93	<b>Dy</b> Dysprosium 162.50	<b>Ho</b> Holmium 164.93	<b>Er</b> Erbium 167.26	<b>Tm</b> Thulium 168.93	<b>Yb</b> Ytterbium 173.05	<b>Lu</b> Lutetium 174.97																
		7	<b>Ac</b> Actinium (227)	<b>Th</b> Thorium 232.04	<b>Pa</b> Protactinium 231.04	<b>U</b> Uranium 238.03	<b>Np</b> Neptunium (237)	<b>Pu</b> Plutonium (244)	<b>Am</b> Americium (243)	<b>Cm</b> Curium (247)	<b>Bk</b> Berkelium (247)	<b>Cf</b> Californium (251)	<b>Es</b> Einsteinium (252)	<b>Fm</b> Fermium (257)	<b>Md</b> Mendelevium (258)	<b>No</b> Nobelium (259)	<b>Lr</b> Lawrencium (266)																

## Conversion of Mass to Energy

Einstein's association of the term  $mc^2$  with a **rest mass energy**  $E_0$  naturally led to speculation about what might be done to *convert* mass into useable energy, since for a *little* mass you get a *lot* of energy!

Let's see *just how much*: in *S.I.* units  $1 \text{ J} \equiv 1 \text{ kg}\cdot\text{m}^2/\text{s}^2$  so a 1 kg mass has a rest mass energy of  $(1 \text{ kg}) \times (2.9979 \times 10^8 \text{ m/s})^2 = 8.9876 \times 10^{16} \text{ J}$  — which is a lot of joules! To get an idea how many, remember that one **watt** is a unit of *power* equal to one joule per second, so a **joule** is the same thing as a **watt-second**. Therefore a device converting *one millionth of a gram* ( $1 \mu\text{g}$ ) of mass to energy *every second* would release approximately *90 megawatts* [millions of watts] of power!

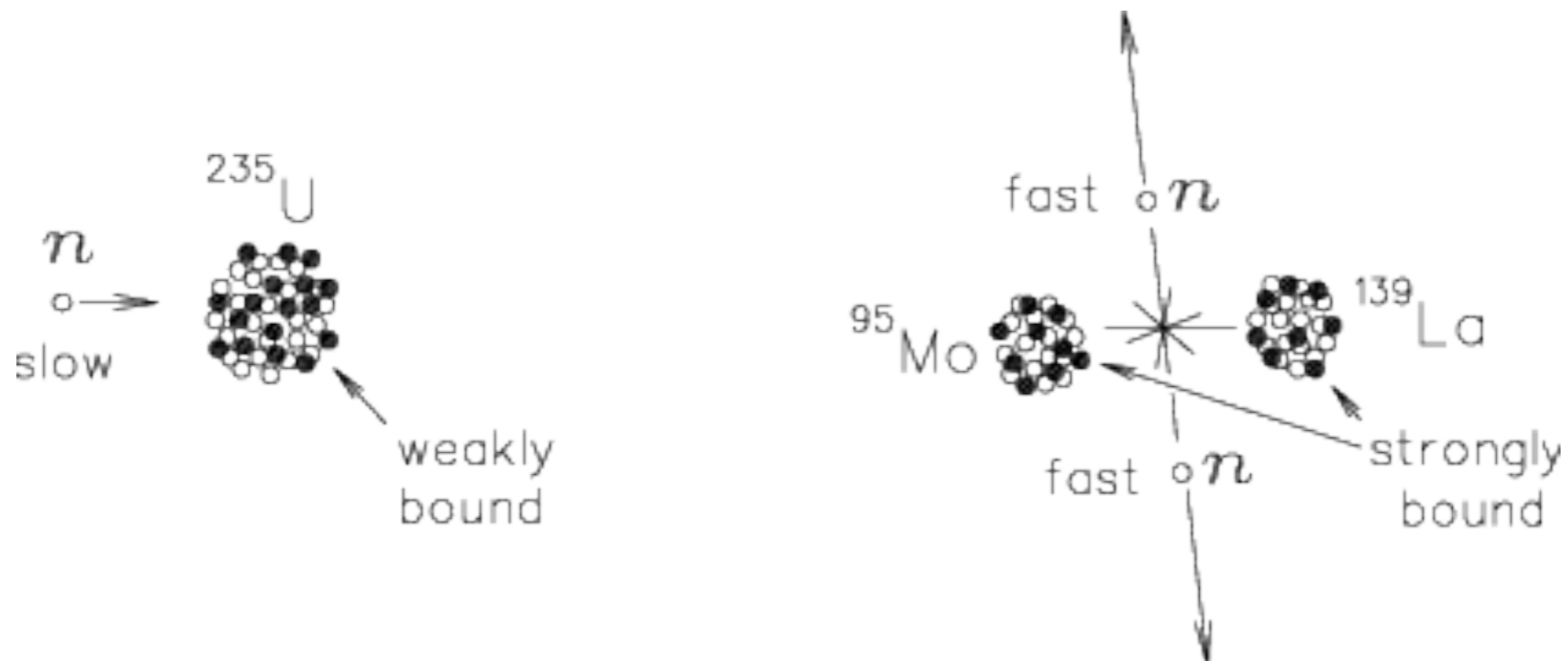
# ***Bad Things to do with Mass-Energy Conversion***

Contrary to popular belief, the first conclusive demonstration of mass-energy conversion was in a controlled nuclear *reactor*.

However, not long after came the more unpleasant manifestation of mass-energy conversion: the fission *bomb*. An unpleasant subject, but one about which it behooves us to be knowledgeable. For this, we need a new energy unit, namely the **kiloton** [kt], referring to the energy released in the explosion of one thousand *tons* of TNT [*TriNitroToluene*], a common chemical high explosive. The basic conversion factor is  $1 \text{ kt} \equiv 1 \text{ trillion calories} = 4.186 \times 10^{12} \text{ J}$ , which means that one **kiloton**'s worth of energy is released in the conversion of 0.04658 grams [46.58 mg] of mass. Thus a **megaton** [equivalent to one *million* tons of TNT or  $10^3$  kt] is released in the conversion of 46.58 grams of mass; and the largest thermonuclear device [bomb] ever detonated, about 100 megatons' worth, converted some 4.658 kg of mass directly into raw energy.

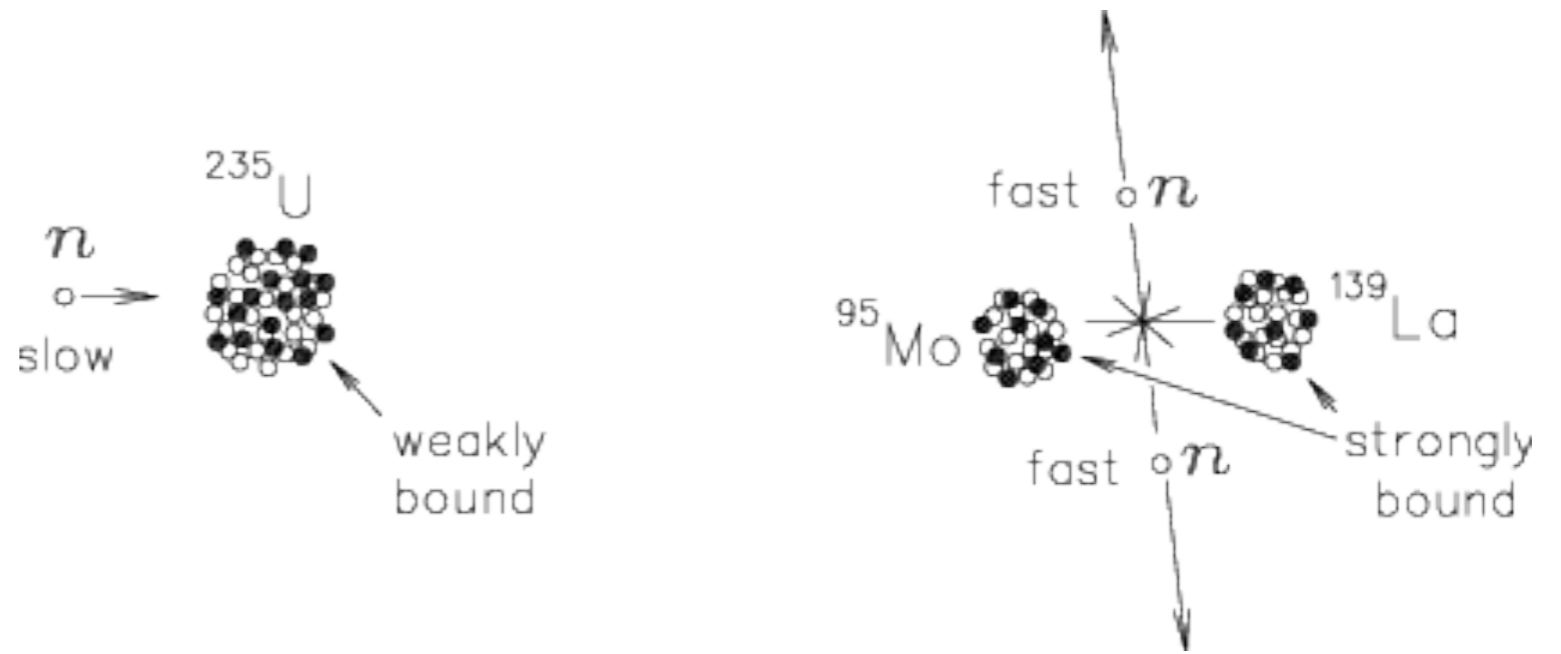
# Nuclear Fission

Where did the energy come from? *What* mass got converted? To answer this question we must look at the processes involved on a sub-microscopic scale. First we must consider the natural tendency for oversized atomic nuclei to spontaneously *split* into smaller components. This process is known as NUCLEAR FISSION and is the energy source for all presently functioning NUCLEAR REACTORS on Earth. [Also for so-called "atomic" bombs.]

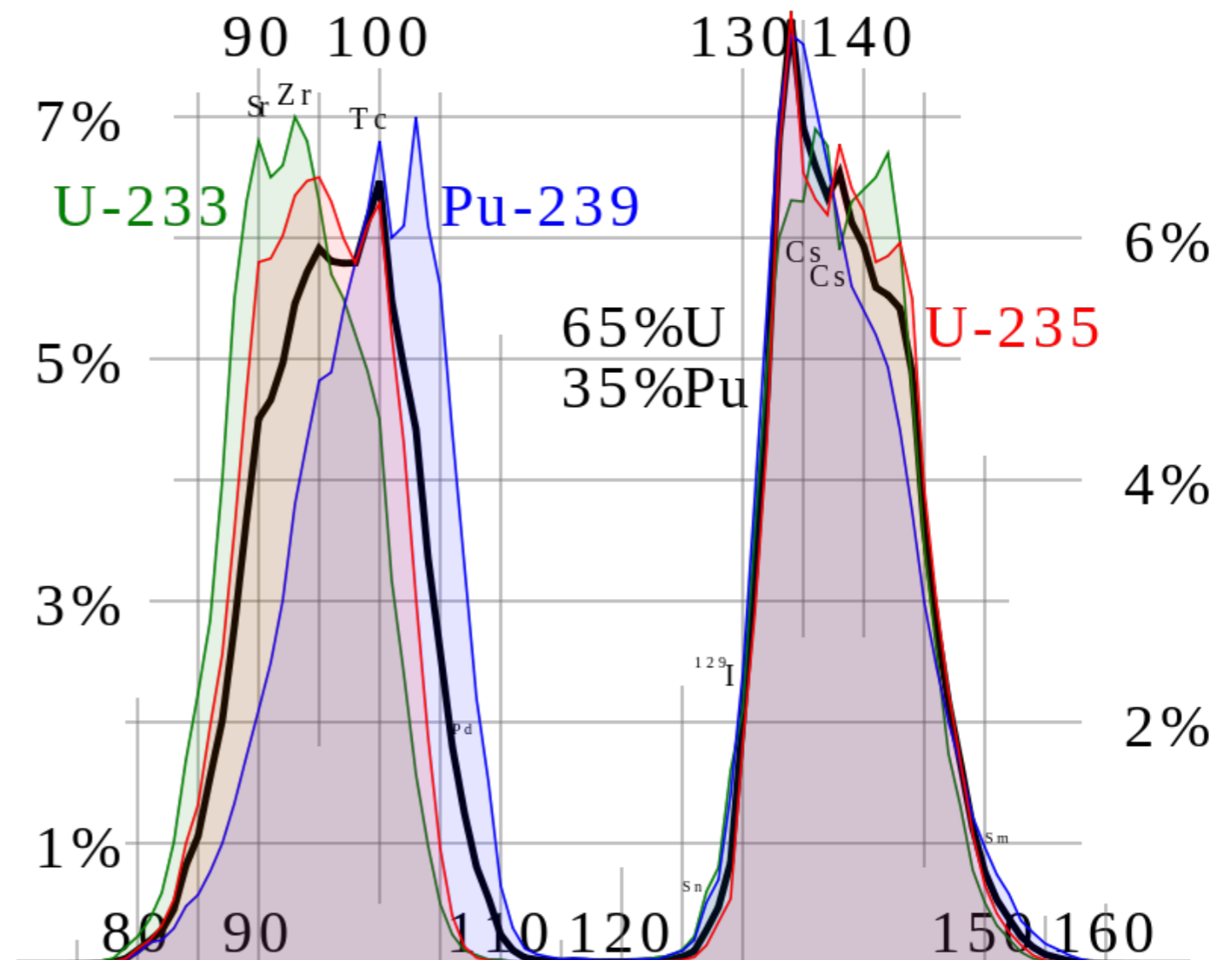


Fission of  $^{235}\text{U}$ : The net mass of the initial neutron plus the  $^{235}\text{U}$  nucleus is  $219,883 \text{ MeV}/c^2$ . The net mass of the fission products (two neutrons, a  $^{95}\text{Mo}$  nucleus and a  $^{139}\text{La}$  nucleus) is  $219,675 \text{ MeV}/c^2$  - smaller because of the stronger *binding* of the Mo and La nuclei. The "missing mass" of  $208 \text{ MeV}/c^2$  (a little less than 0.1% of the original mass) goes into the *kinetic energy* of the *fragments* (mainly the *neutrons*), which of course adds up to 208 MeV.

# FISSION



The example shown above has  $^{235}\text{U}$  splitting into  $^{95}\text{Mo}$  and  $^{139}\text{La}$  along with *two* fast neutrons, but there are *many other ways* that  $^{235}\text{U}$  can fission. One of its "daughters" usually has 90-100 nucleons, while the other has 130-140. And sometimes there are more than two neutrons emitted. This, plus the long lifetime of  $^{235}\text{U}$  and its relatively innocuous alpha-decay, make it a "good" candidate for a **bomb**.

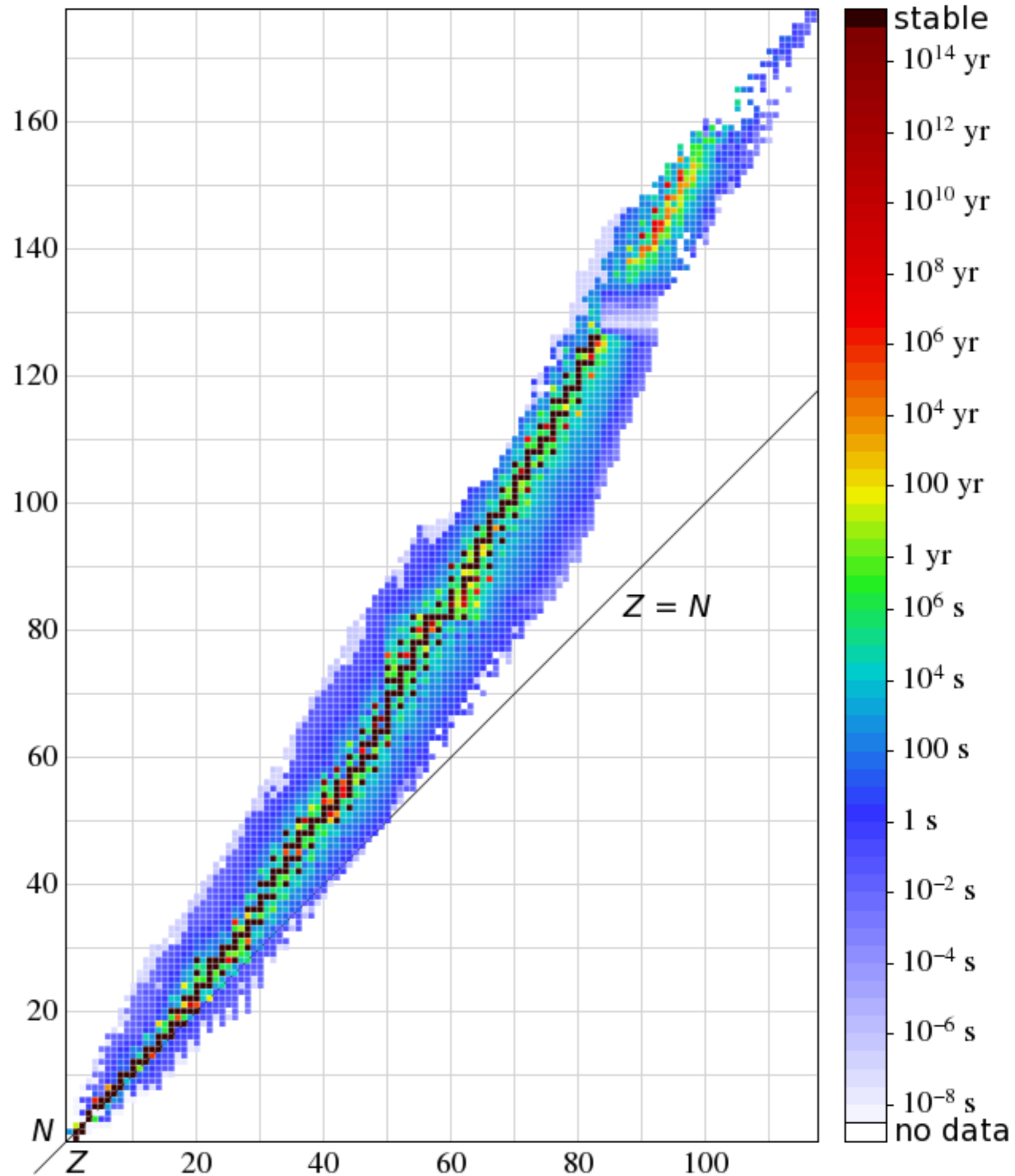




# ISOTOPES and their (color-coded) Half-Lives

**$Z$  = # of protons**

**$N$  = # of neutrons**

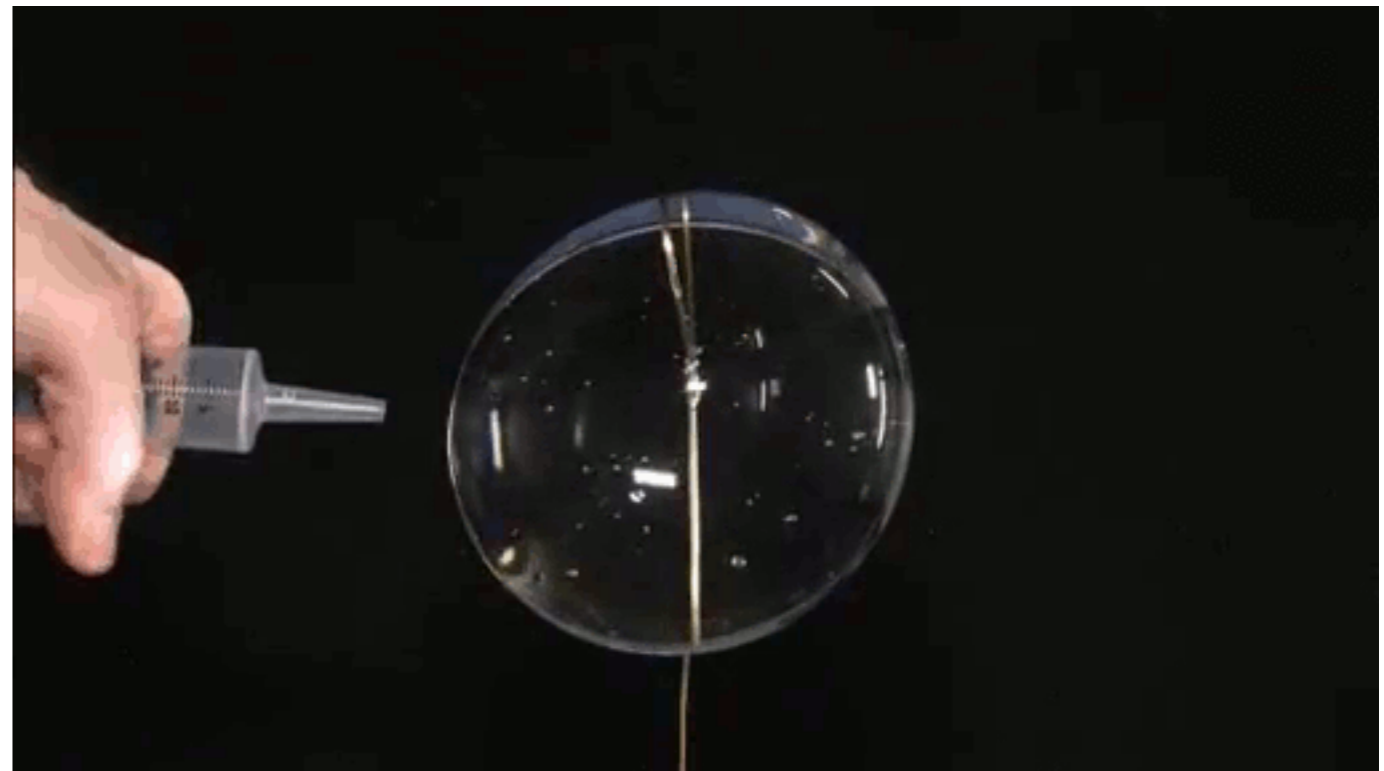


# Liquid Drop Model



Surface tension keeps *small* water droplets perfectly spherical, but **big** water droplets are *unstable* because the attractive force between molecules is *short-range*.

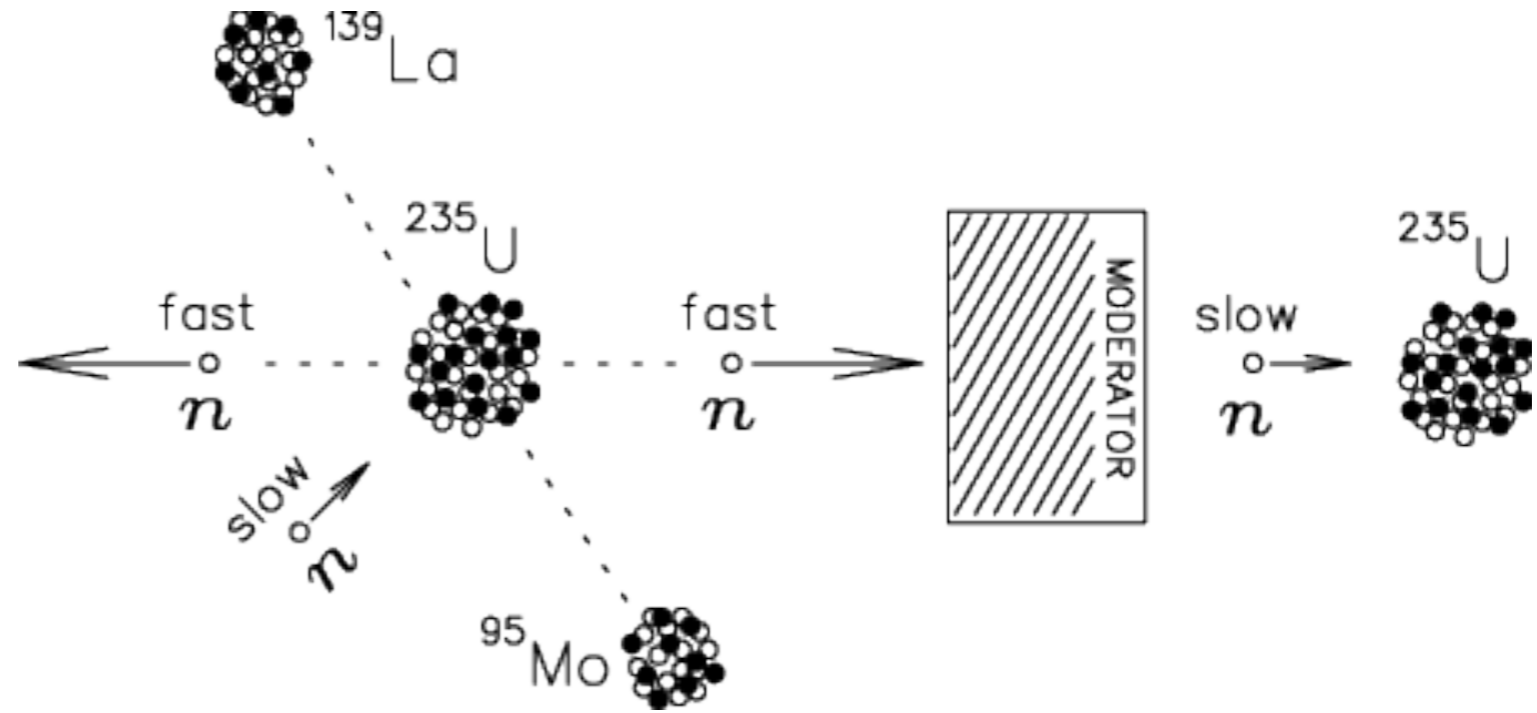
**Nuclei** are similar: *small* ones are *stable*, but **big** ones jiggle and sometimes *split* into smaller nuclei, spitting out neutrons, for the same reason: the **strong attraction** between nucleons is *short range*, while the electrostatic *repulsion* between *protons* is *long range*.



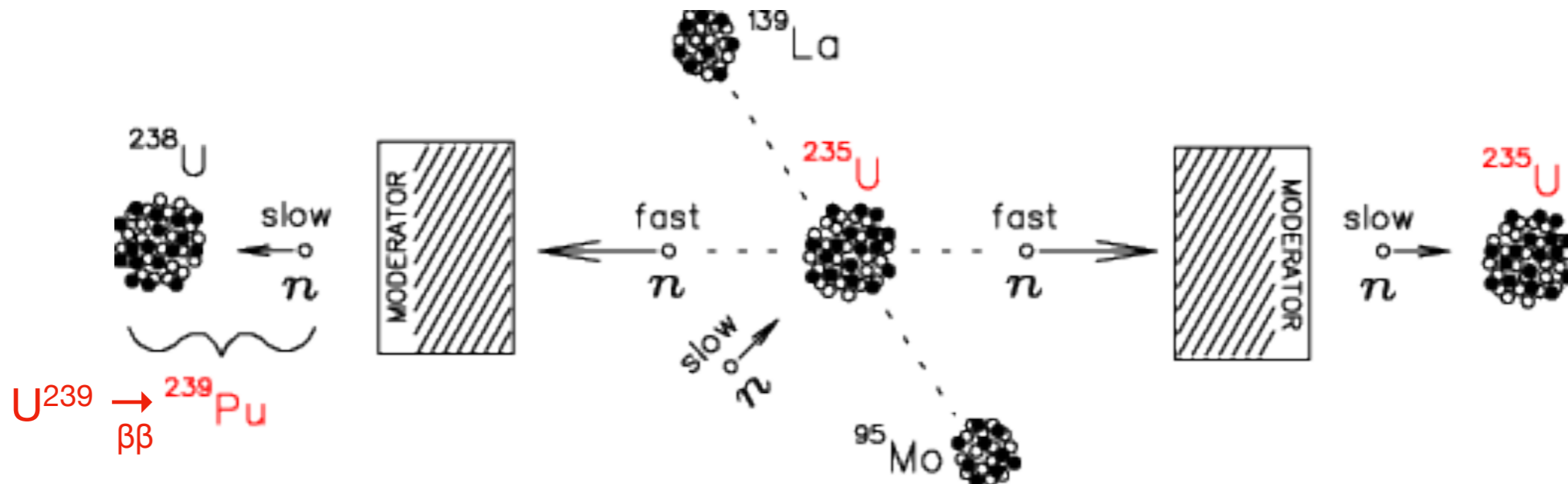


# Chain Reactions:

- $^{235}\text{U}$  by itself:



- Breeding *fissile*  $^{239}\text{Pu}$  from *fertile*  $^{238}\text{U}$  with *neutrons* from  $^{235}\text{U}$ :



# How (e.g.) $^{235}\text{U}$ Bombs Work

1. Sub-critical chunks of  $^{235}\text{U}$  are blasted together by conventional explosives, producing a temporary blob of *extremely* compressed  $^{235}\text{U}$ .
2. A stray neutron is captured by  $^{235}\text{U}$ , causing it to fission into two lighter nuclei and *several* neutrons.
3. Those neutrons slow down in the hyperdense metal and capture on other  $^{235}\text{U}$  nuclei, causing *them* to fission...

CHAIN REACTION!

# How *(some)* “H-Bombs” Work

1. Set off a small (*e.g.*)  $^{235}\text{U}$  **fission** bomb to produce extreme pressure and temperature.
2. This triggers the **fusion** of (*e.g.*) deuterium & tritium, releasing huge amounts of energy, mainly in the form of *fast neutrons*.
3. *Fast* neutrons hit surrounding  $^{238}\text{U}$  nuclei, causing *them* to **fission**...
4. Add as much  $^{238}\text{U}$  as you like. It's a **fission** bomb!

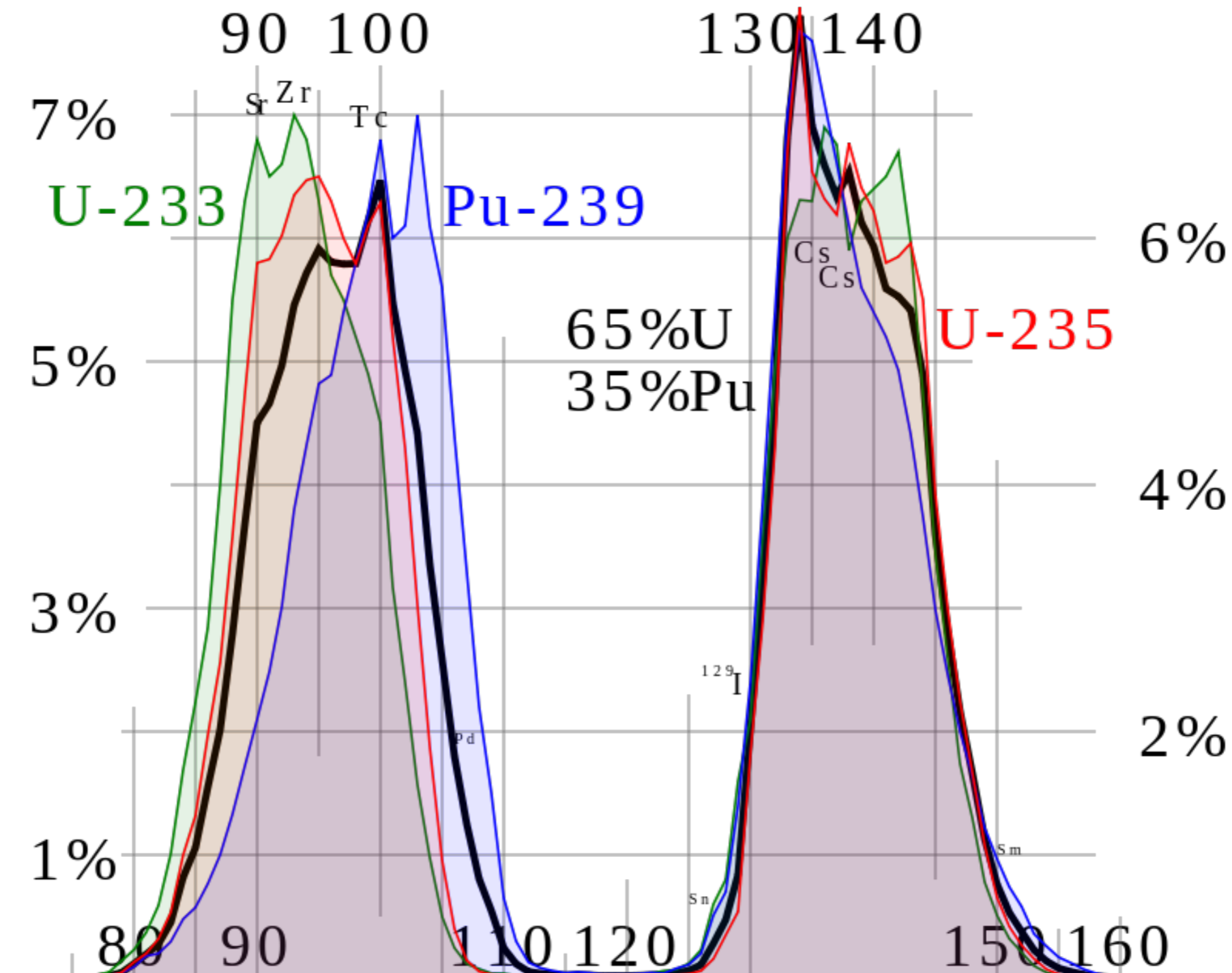
# How **Reactors** Work

- Establish a **controlled** chain reaction in fission “fuel” using a combination of **moderators** (to slow down the fast fission neutrons so they can be captured and cause additional fissions) and neutron-absorbing “**control rods**” to decrease the slow neutron flux.
- **Heat** results from the energy deposited. Use it to drive a thermodynamic **heat engine** (e.g. a steam turbine). The hotter, the higher the efficiency.

$$(\text{Carnot efficiency} = 1 - T_{\text{cold}}/T_{\text{hot}})$$

- That’s it. There are almost no exceptions *so far*.

# Fission Products



# Decay of Fission Products

Wikipedia: “The radioactivity in the fission product mixture is **[initially]** mostly *short-lived* isotopes such as  $^{131}\text{I}$  and  $^{140}\text{Ba}$ ; after about four months  $^{141}\text{Ce}$ ,  $^{95}\text{Zr}/^{95}\text{Nb}$  and  $^{89}\text{Sr}$  take the largest share, while after about two or three years the largest share is taken by  $^{144}\text{Ce}/^{144}\text{Pr}$ ,  $^{106}\text{Ru}/^{106}\text{Rh}$  and  $^{147}\text{Pm}$ .”

## Medium-lived Daughters

Prop:	$t_{1/2}$	Yield	Q *	$\beta\gamma$ *
Unit:	(a)	(%)	(keV)	
$^{155}\text{Eu}$	4.76	0.0803	252	$\beta\gamma$
$^{85}\text{Kr}$	10.76	0.2180	687	$\beta\gamma$
$^{113\text{m}}\text{Cd}$	14.1	0.0008	316	$\beta$
$^{90}\text{Sr}$	28.9	4.505	2826	$\beta$
$^{137}\text{Cs}$	30.23	6.337	1176	$\beta\gamma$
$^{121\text{m}}\text{Sn}$	43.9	0.00005	390	$\beta\gamma$
$^{151}\text{Sm}$	88.8	0.5314	77	$\beta$

## Long-lived Daughters

Nuclide	$t_{1/2}$	Yield	Decay energy <sup>[a 1]</sup>	Decay mode
	(Ma)	(%) <sup>[a 2]</sup>	(keV)	
$^{99}\text{Tc}$	0.211	6.1385	294	$\beta$
$^{126}\text{Sn}$	0.230	0.1084	4050 <sup>[a 3]</sup>	$\beta\gamma$
$^{79}\text{Se}$	0.327	0.0447	151	$\beta$
$^{93}\text{Zr}$	1.53	5.4575	91	$\beta\gamma$
$^{135}\text{Cs}$	2.3	6.9110 <sup>[a 4]</sup>	269	$\beta$
$^{107}\text{Pd}$	6.5	1.2499	33	$\beta$
$^{129}\text{I}$	15.7	0.8410	194	$\beta\gamma$



# Health Concerns

Isotope	Radiation	Half-life	GI absorption	Notes
Strontium-90/yttrium-90	$\beta$	28 years	30%	
Caesium-137	$\beta, \gamma$	30 years	100%	
Promethium-147	$\beta$	2.6 years	0.01%	
Cerium-144	$\beta, \gamma$	285 days	0.01%	
Ruthenium-106/rhodium-106	$\beta, \gamma$	1.0 years	0.03%	
Zirconium-95	$\beta, \gamma$	65 days	0.01%	
Strontium-89	$\beta$	51 days	30%	
Ruthenium-103	$\beta, \gamma$	39.7 days	0.03%	
Niobium-95	$\beta, \gamma$	35 days	0.01%	
Cerium-141	$\beta, \gamma$	33 days	0.01%	
Barium-140/lanthanum-140	$\beta, \gamma$	12.8 days	5%	
Iodine-131	$\beta, \gamma$	8.05 days	100%	
Tritium	$\beta$	12.3 years	100%	[a]