### Diffraction as "∞-Slit Interference"

If  $I_0$  is the intensity at the central maximum,  $A_0 = \sqrt{I_0}$  is the amplitude at the central maximum. By dividing up the slit into N infinitesimal "pseudoslits", each contributing an amplitude  $A_1 = A_0/N$ , and letting  $N \to \infty$ , we get

$$A = A_0 (\sin \alpha) / \alpha$$
 or  $I = I_0 [(\sin \alpha) / \alpha]^2$ 

where  $\alpha \equiv \pi (a \sin \theta) / \lambda$  and a = the width of the slit.

This diffraction pattern has its <u>first minimum</u> where  $\alpha = \pi$  or

$$\frac{a}{3}\sin\frac{a}{3}=\lambda$$
.

The maxima of  $I(\alpha)$  are found where the slope of  $(\sin \alpha)/\alpha$  is zero: setting the derivative equal to zero gives (after some algebra)  $\tan \alpha = \alpha$ which is a transcendental equation best solved graphically.

#### Maxima of Diffraction Pattern

The maxima of  $I(\alpha)$  occur where  $\tan \alpha = \alpha$ , a transcendental equation best solved graphically:



# **N-Slit Gratings**



Be careful not to confuse  $9_{1i}$  with  $9_{1d}$ ! The formulae look alike but...

## **Circular Apertures**

We have been talking about "slits" as if all diffraction problems were onedimensional. In reality, the most common type is circular, such as telescopes, laser cannons and the pupil of your eye. The following handwaving logic is not a proof, but a plausibility argument:

The narrower the slit, the wider the diffraction pattern. (Look at the formula for the first minimum!) Picture a circular aperture as a square aperture with the "corners chopped off": on average, it is narrower than the original square whose side was equal to the circle's diameter. Thus you would expect it to produce a wider diffraction pattern. Indeed it does! The numerical difference is a factor of 1.22: instead of  $a \sin \theta_1 = \lambda$  we have

 $a \sin \theta_1 = 1.22 \lambda$ .

### Resolution

Diffraction is "reversible" in the sense that two light sources **inside** a telescope shining out through an aperture would produce overlapping patterns at a distant detector for the same conditions that two point sources at the **detector's** location would no longer be **resolvable** by the telescope's optics.

Thus the criterion for "resolvability" of two distant stars (for example) by a telescope of diameter a is that their angular separation be greater than the angle  $g_{1d}$  given by  $a \sin g_1 = 1.22 \lambda$  between each one's central maximum and the first minimum of its diffraction pattern. This "Rayleigh criterion" also applies for microscopes, the pupil of the eye, etc.

# Dispersion

Since the angular pattern of interference and diffraction from a grating explicitly depends on the wavelength  $\lambda$ , it follows that light of different wavelengths will be "bent" by different angles. If we consider the  $m^{\text{th}}$  Principal Maximum, for which  $d \sin \vartheta_m = m \lambda$ , and take the derivative of both sides with respect to  $\lambda$ , we get  $d \cos \vartheta_m (d \vartheta_m / d \lambda) = m$  or

$$d\vartheta_m/d\lambda = m/d\cos\vartheta_m$$
 giving  $d\vartheta_m/d\lambda \approx m/d$  for small angles.

Thus if the purpose of our grating is to "resolve" different colours (known as "dispersion") then we want to have the smallest possible slit spacing d and the largest possible "order" m.

 $d\vartheta_m/d\lambda \equiv D$  is known as "the Dispersion" of a grating.

# **Resolving Power**

Again assuming we are using a grating to measure  $\lambda$ , <u>how close together</u> ( $\Delta\lambda$ ) <u>can two wavelengths be</u>  $(\lambda' = \lambda + \Delta\lambda)$  <u>and still be resolved</u>? That is, if  $\lambda$ has an  $m^{\text{th}}$  order Principal Maximum (PM<sub>m</sub>) at  $\mathcal{G}_m$  then  $\lambda'$  has <u>its</u> PM<sub>m</sub> at  $\mathcal{G}_m' = \mathcal{G}_m + \Delta \mathcal{G}_m$  right on top of the <u>1<sup>st</sup> minimum beyond</u>  $\mathcal{G}_m$  for  $\lambda$ .



We know that  $d \sin \vartheta_m = m\lambda$ ,  $d \sin \vartheta_m' = m\lambda'$ and  $d \sin (\vartheta_m + \Delta \vartheta_m) = (m + 1/N) \lambda$  so that the extra path length difference between adjacent slits,  $\lambda/N$ , ensures a phasor diagram that closes on itself in N phasors. So we must have  $m\lambda' = (m + 1/N) \lambda$  or  $\lambda + \Delta\lambda = \lambda + \lambda/mN$  or  $\lambda/\Delta\lambda = mN \equiv R$ ,

the "Resolving power" of the grating.